

The Infrared Physics in QED and cosmology

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Introduction: infrared triangle

Infrared triangle

Strominger (2013, 2017)

Strominger, and Zhiboedov (2014)

- Infrared dynamics is represented by the three equivalent subjects.
- This equivalence is expected to be valid in various theories with massless particles (e.g. **linearized GR**, **QED**, **YM theory**, cosmology,...).

Mostly, gauge theory in asymptotically flat spacetime

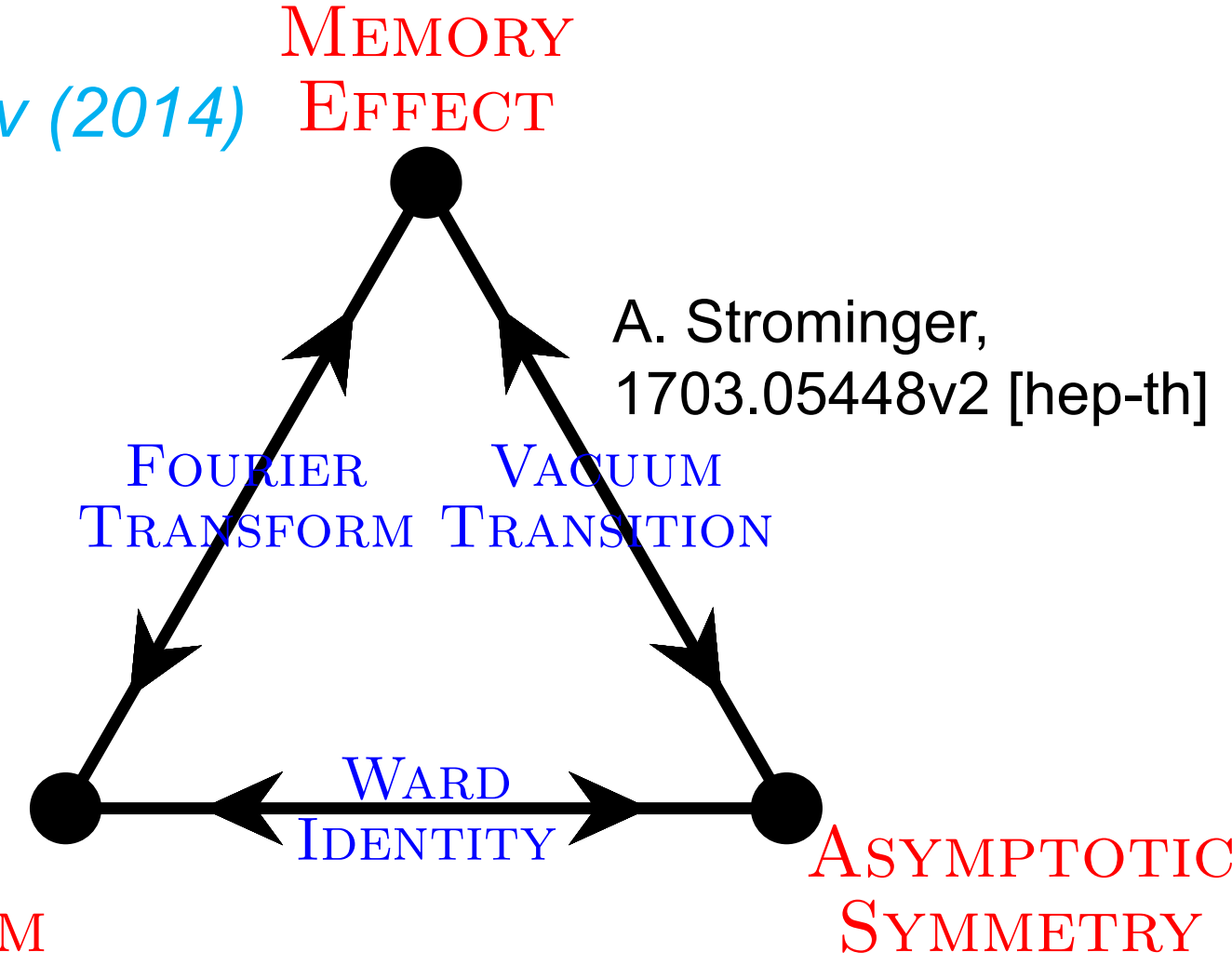


Fig. 1 The infrared triangle

Goals of this work

- to understand the locality condition, which ensure the effective description of primordial perturbation.
- to extend IR triangle to $\omega \neq 0$ modes

Special features of cosmology (inflation)

- nearly de Sitter spacetime (not asymptotically flat)
- extended to $k \neq 0$ modes

Weinberg's soft theorem and IR cancellation in QED

Weinberg (1965)

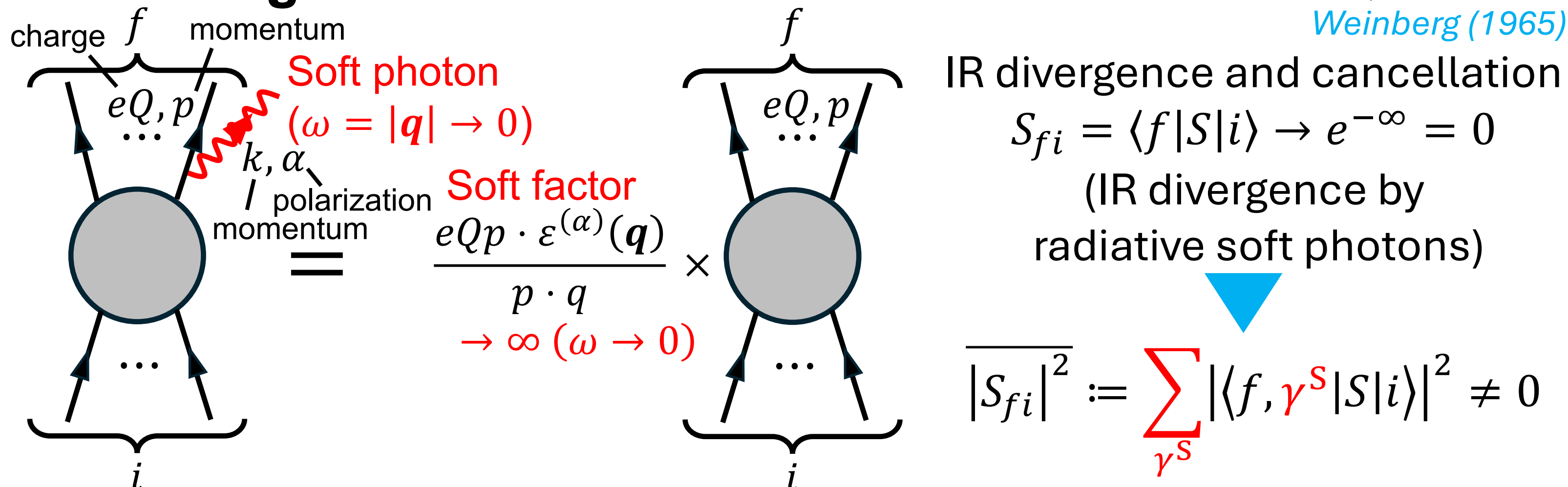


Fig. 2 Weinberg's soft theorem for photon

Asymptotic symmetry in QED

Strominger (2013, 2017)

He, Mitra, Porfyriadis, Strominger (2014)

Large gauge transformation (LGT) (= asymptotic symmetry in QED)

$$\begin{cases} A_\mu(x) \rightarrow A_\mu(x) + \partial_\mu \lambda(x) \\ \psi(x) \rightarrow e^{iQ\lambda(x)}\psi(x) \end{cases} \quad \text{with} \quad \lambda(x) \xrightarrow[r \rightarrow \infty (t \pm r: \text{fixed})]{\text{asymptotic limit}} \varepsilon(\hat{x}) \neq 0$$

Generator of LGT (out region)

("(0)" means the leading order in $1/r$ expansion.)

$$Q_\varepsilon^+ := \frac{1}{e^2} \int dud^2\hat{x} \left[-\partial_A \varepsilon(\hat{x}) F_u^{(0)A}(u, \hat{x}) + e^2 \varepsilon(\hat{x}) j_u^{(0)}(u, \hat{x}) \right] \quad (F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu, u = t - r)$$

Q_ε^+ includes only $\omega = 0$ photons by integral with respect to u .

Ward identity (conservation of the LGT charge)

$$\langle f | (Q_\varepsilon^+ S - S Q_\varepsilon^-) | i \rangle = 0$$

$$\varepsilon(z, \bar{z}) = \text{const.} \Rightarrow \text{total electric charge conservation}$$

$$\varepsilon(z, \bar{z}) = \frac{1}{z - w} \Rightarrow \text{Weinberg's soft photon theorem} \quad (z, w \in \mathbb{C} \text{ are coordinates on } S^2 \text{ (stereographic projection)})$$

Faddeev-Kulish's dressed state and IR cancellation

$$|\alpha; \text{FK}\rangle := e^{iR} |\alpha\rangle, \quad (\alpha = i, f)$$

R generates cloud of soft photons

$$Q_\varepsilon^- |i; \text{FK}\rangle = 0, \quad \langle f; \text{FK} | Q_\varepsilon^+ = 0$$

asymptotic symmetry

$$S_{fi} = \langle f | S | i \rangle \propto e^{-\infty} = 0$$

(IR divergence by radiative soft photons)

$$S_{fi}^{\text{FK}} = \langle f; \text{FK} | S | i; \text{FK} \rangle \neq 0 \quad (\text{IR cancellation})$$

Cosmology and extension of infrared physics in QED

Infrared physics in cosmology (inflation)

Tanaka & Urakawa (2017, 2026)

Dilatation (LGT in cosmology)

$$\begin{cases} x^i \rightarrow x_\lambda^i = e^\lambda x^i \\ \zeta(t, x) \rightarrow \zeta_\lambda(t, x) = \zeta(t, e^{-\lambda} x) - \lambda \end{cases} \quad (\zeta: \text{curvature perturbation } (ds^2 \supset e^{2\zeta} dx^2))$$

Generator

$$Q[\lambda] = a^3(t) \int d^3x \Delta_\lambda \zeta(t, x) \pi_\zeta(t, x) \quad (a(t): \text{scale factor}, \Delta_\lambda \zeta: \text{linear-order deviation})$$

Dilatation invariance (large gauge symmetry)
 $Q[\lambda] |\Psi\rangle = 0$
($|\Psi\rangle$: state of universe)

&

Locality condition (inflation)

$$\langle \zeta^c | \Psi \rangle = |\psi(\zeta^c)| \prod_a |\Psi_a; \zeta_a^c\rangle$$

$$(\zeta_a | \zeta_a^c) = \zeta_a^c | \zeta_a^c), \quad |\zeta^c\rangle = \prod_a |\zeta_a^c\rangle, \quad \zeta_a^c: \text{average of } \zeta \text{ over the patch } a)$$

Soft theorem for $k \neq 0$

(generalized consistency relation)

$$\langle \Psi | \zeta_k^S \mathcal{O}_{k'} | \Psi \rangle \propto \langle \Psi | \zeta_k^S \zeta_{k'}^S | \Psi \rangle \langle \Psi | \mathcal{O}_{-k-k'} | \Psi \rangle$$

Fig. 3 Separate universe in cosmology

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LGT invariance in $\partial_z \phi$ -representation (Q_H^+ : fermion part of the LGT generator)

$$\frac{\partial \mathcal{N}}{\partial (\partial_A \phi^c)} = 0, \quad \left(iQ_H^+ - \partial_A \varepsilon \cdot \frac{\partial}{\partial (\partial_A \phi^c)} \right) |f; \partial_z \phi^c\rangle = 0$$

LGT invariance is rewritten in the similar form as that in cosmology.

Separation of asymptotic region in QED

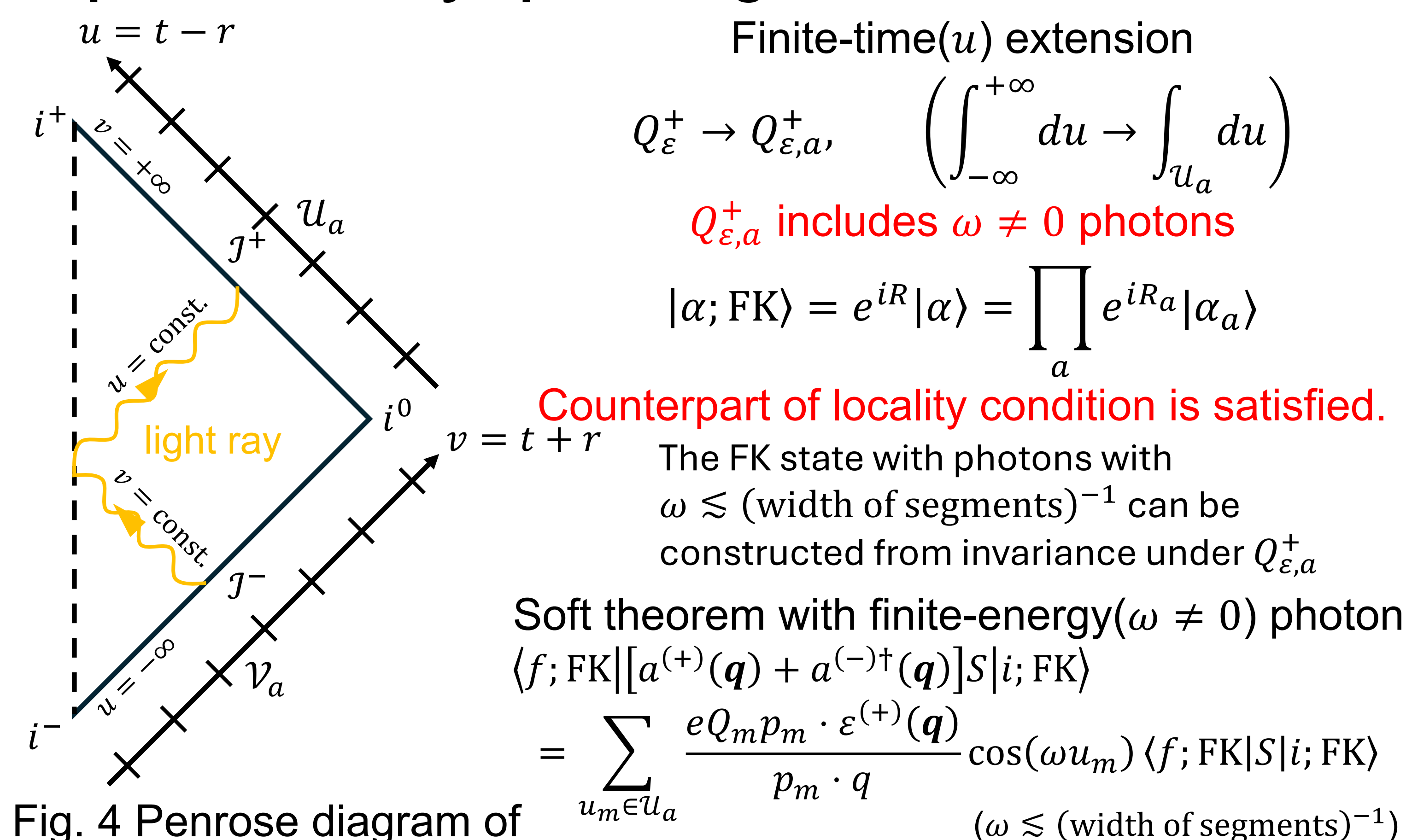


Fig. 4 Penrose diagram of

Minkowski spacetime ($r \geq 0$)

Similarities and differences

Similar points

- The correspondence of quantities is made (Table. 1).
- LGT invariance for hard state can be written in the similar form.

Different points

- The magnitude of the wave function is constant or flat in one direction,
- and so on.

Summary and prospects

Summary

- LGT invariance is rewritten in the similar form to that in cosmology.
- The magnitude of wave function is constant in QED.
- Soft theorem with $\omega \neq 0$ is obtained by separating $J^\pm$.
- Separation of Q_ε^+ is needed to obtain the soft theorem for $\omega \neq 0$.
- The counterpart of the locality condition is satisfied for FK state in QED.

Prospects

- electromagnetic memory in finite time? How can it be detected?
- physical meaning of the finite-energy formulation in QED?

Conjugate coherent state in QED

TK, Tanaka, & Urakawa (in progress)

$$N_z^+(z, \bar{z}) |f; \text{FK}\rangle := N_z^{\text{out}}(z, \bar{z}) |f; \text{FK}\rangle, \quad \text{Table. 1 Corresponding quantities}$$

$$N_z^+(z, \bar{z}) = \int du F_{uz}^{(0)}(u, z, \bar{z}) \quad \text{gauge invariant}$$

ζ is large gauge dependent.

$$\zeta(t, x) \Leftrightarrow A_z^{(0)}(u, z, \bar{z}) \quad (\text{not } N_z^+)$$

For unseparated case,

$$\bar{\zeta}(t) \text{ (spatial average)} \Leftrightarrow \partial_z \phi(z, \bar{z}) \text{ (average w.r.t. } u)$$

$$\partial_z \phi \text{ is conjugate to } N_z^\pm: [N_z^\pm(z, \bar{z}), \partial_z \phi(z', \bar{z}')] := -\frac{ie^2}{2} \delta^2(z - z')$$

introduce conjugate coherent state: $\partial_z \phi | \partial_z \phi^c \rangle = \partial_z \phi^c | \partial_z \phi^c \rangle$

$$|f; \text{FK}\rangle = \int \mathcal{D} \partial_z \phi^c | \partial_z \phi^c \rangle \langle \partial_z \phi^c | f; \text{FK} \rangle, \quad \begin{cases} [\hat{p}, \hat{x}] = -i \\ \hat{x} |x\rangle = x |x\rangle, \quad \hat{p} |p\rangle = p |p\rangle \\ \langle x | p \rangle = C e^{ipx} \end{cases}$$

$$\langle \partial_z \phi^c | f; \text{FK} \rangle = \mathcal{N} e^{2ie^{-2} N_z^{\text{out}} A \cdot \partial_A \phi^c} |f\rangle$$