### Hawking-Page and entanglement phase transition in 2d CFT on curved backgrounds Akihiro Miyata (KITS,UCAS)

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- We consider a two-dimensional holographic CFT defined on curved "Hawking-Page like" phase transition happens.
- (effective) temperature.

backgrounds, and by varying the CFT background metric we observe that

• We also consider the gravity dual of the thermal state of the CFT for high

## Contents of this talk

- Introduction 1.
  - AdS/CFT
  - Hawking-Page transition
- 2. Global and Local properties of inhomogeneous thermal state
- 3. Gravity dual to the CFT

Summary and Future directions

Inhomogeneous Hamiltonian and Hamiltonian on a curved background

### 1. Introduction

#### AdS/CFT and Holographic CFTs AdS/CFT correspondence [Maldacena '97,...]

- - Quantum gravity on (d + 1)-dim. Anti-de Sitter spacetime (AdS)
    - = d-dim. Conformal field theory (CFT)
- CFTs, having semi-classical gravity ( $G_N \ll 1$ ) duals, are called Holographic CFTs.
- Holographic CFTs are characterized by the following conditions [EI-Showk Papadodimas '11]
  - It has large central charge i.e. many degrees of freedom. 1.
  - It has a small number of operators of low conformal dimension. 2.
  - 3. The correlators of the low-lying operators factorize.
- For simplicity, focus on AdS<sub>3</sub>/CFT<sub>2</sub>.



CFT

# Hawking-Page transition

- - Low temperature  $L/\beta < 1 \rightarrow$  thermal AdS
  - High temperature  $L/\beta > 1 \rightarrow AdS$  (BTZ) black holes
- deconfinement phase transition [Witten '98, Maldacena '98, ...]: Thermal AdS  $\rightarrow$  confined phase, BTZ  $\rightarrow$  de-confined phase

• In AdS semi-classical gravity  $G_N \ll 1$  ( $\leftrightarrow c \gg 1$ ), dominant saddle geometries are given by either thermal AdS or AdS black holes depending on the temperature  $T = \beta^{-1}$  and system size L; there is a phase transition between them, ie., Hawking-Page transition [Hawking-Page '83,...],



Thermal AdS

In CFT, this phase transition corresponds to a first-order confinement-



### **Deformation of the CFT Hamiltonian**

• Start with the deformation of the the system on the spatial circle with circumstance L

$$H_0 = \int_0^L rac{dx}{2\pi} (T(x) + \overline{T}(x)) = rac{2\pi}{L} igg[ L_0^z + \overline{L}_0^{\overline{z}} - rac{c}{12} igg] \qquad (z, ar{z}) = \Big( e^{rac{2\pi (ix + \mathcal{T})}{L}}, e^{rac{2\pi (ix + \mathcal{T})}{L}} igg]$$

$$egin{aligned} &
ightarrow H_{q- ext{ M\"obius }} = \int_0^L rac{dx}{2\pi} \Big[ 1- anh 2 heta \Big( 1-2 \sin^2 \Big( rac{q\pi x}{L} \Big) \Big) \Big] (T(x)+\overline{T}(x)) \ &= rac{2\pi}{L} \Big[ L_0^z + \overline{L}_0^z - rac{ anh 2 heta \Big( L_q^z + \overline{L}_q^z + L_{-q}^z + \overline{L}_{-q}^z \Big) - rac{c}{12} \Big] \end{aligned}$$

- The spatial inhomogeneity is introduced by the enveloping function

$$f(x, heta) = 1 - anh 2 heta \Big( 1 - 2 \sin^2 \Big( rac{q \pi x}{L} \Big) \Big) = 1 - anh 2 heta \cdot \cos igg( rac{2q \pi x}{L} igg)$$

Sine-square deformed (q-SSD) Hamiltonian.

q : Positive integer,  $\theta$  : non-negative real parameter

• At  $\theta = 0$ , the Hamiltonian reduces to the uniform one, and at the large- $\theta$ , it becomes the q-deformed



# CFT on a curved background

 The inhomogeneous Hamiltonian is equivalent to the uniform Hamiltonian on the curved background with the metric

$$egin{aligned} H_{q-\, ext{M\"obius}} &= \int_{0}^{L} rac{dx}{2\pi} \sqrt{-\det g(x)} (T(x) + ar{T}(x)), \ ds^2 &= g_{ab}(x) dx^a dx^b = -f(x, heta)^2 dt^2 + dx^2 \end{aligned}$$

The CFT background has the Ricci curvature

$$R = -rac{2\partial_x^2 f(x, heta)}{f(x, heta)} = rac{8\pi^2 q^2 anh(2 heta) \cos\left(rac{2\pi q x}{L}
ight)}{L^2 igl( anh(2 heta) \cos\left(rac{2\pi q x}{L}
ight) - 1igr)}$$
 $f(x, heta) = 1 - anh 2 heta \cdot \cos\left(rac{2q\pi x}{L}
ight)$ 



#### • The CFT background has the Ricci curvature

$$R = -rac{2\partial_x^2 f(x, heta)}{f(x, heta)} = rac{8\pi^2 q^2 anh(2 heta) \cos\left(rac{2\pi qx}{L}
ight)}{L^2 \Big( anh(2 heta) \cos\left(rac{2\pi qx}{L}
ight) - 1\Big)},$$
 $f(x, heta) = 1 - anh 2 heta \cdot \cos\left(rac{2q\pi x}{L}
ight)$ 



 $L^2$ R







#### Thermal state for the Hamiltonian Consider the thermal state of using the Hamiltonian, not having the form of

- the uniform one
- By the conformal maps

$$ig(z,ar{z}ig)=ig(e^{rac{2\pi(ix+\mathcal{T})}{L}},e^{rac{2\pi(-ix+\mathcal{T})}{L}}ig), \chi=ig(rac{\cosh( heta)z^q-\sinh( heta)}{\cosh( heta)-\sinh( heta)z^q}ig)^{rac{1}{q}}, ar{\chi}=ig(rac{\cosh( heta)ar{z}^q-\sinh( heta)}{\cosh( heta)-\sinh( heta)ar{z}^q}ig)^{rac{1}{q}},$$

the Hamiltonian is mapped to that having the standard form like the uniform Hamiltonian,

$$egin{aligned} H_{ ext{q-M\"obius}} &= rac{2\pi}{L_{ ext{eff}}} \Big[ L_0^\chi + ar{L}_0^{ar{\chi}} - rac{c}{12} \Big] + rac{2\pi cq^2}{12L_{ ext{eff}}} - rac{2\pi cq^2}{12L} & \left( H_0 = rac{2\pi}{L} \Big[ L_0^z + ar{L}_0^{ar{z}} - rac{c}{12} \Big] 
ight) \ & (L_{ ext{eff}} = L \cosh 2 heta : ext{effective system size}) \ & 
ho = e^{-eta H_{q-M\"obius}} / ext{tr} \, e^{-eta H_{q-M\"obius}} = \exp \Bigg( rac{-2\pieta \Big( L_0^\chi + \overline{L}_0^{ar{\chi}} \Big) }{L_{ ext{eff}}} \Bigg) / ext{tr} \exp \Bigg( rac{-2\pieta \Big( L_0^\chi + \overline{L}_0^\chi \Big) }{L_{ ext{eff}}} \Bigg) / ext{tr} \exp \Bigg( rac{-2\pieta \Big( L_0^\chi + \overline{L}_0^\chi \Big) }{L_{ ext{eff}}} \Bigg) \end{aligned}$$

 $ho=e^{-eta H_{q-\,\mathrm{M\"obius}}}\,/\mathrm{tr}\,e^{-eta H_{q-\,\mathrm{M\"obius}}}$ 

- Thus, by using the  $(\chi, \bar{\chi})$  coordinates,
  - the thermal state for the inhomogeneous CFT Hamiltonian with size L
    - (= CFT Hamiltonian on the curved spacetime)
- = the thermal state for the uniform CFT Hamiltonian with replacing  $\frac{\beta}{L}$  with  $\frac{\beta}{L\cosh 2\theta}$ 
  - $(= CFT Hamiltonian on a (conformally) flat spacetime with size L cosh 2\theta)$
- In the coordinates, we can use usual techniques for the standard thermal state



## Main result

- We studied thermal and entanglement properties of the inhomogeneous thermal state by changing the parameter  $\theta$  controlling the CFT background metric in a two-dimensional holographic CFT.
- We found that, by changing  $\theta$  with fixing  $L/\beta < 1$ , we observed that Hawking-Page like phase transition happens.
- We also consider the gravity dual for the inhomogeneous thermal state for  $L \cosh(2\theta)/\beta > 1$ .

# Contents of this talk

- 1. Introduction  $\checkmark$
- 2. Global and Local properties of inhomogeneous thermal state
  - Thermal entropy
  - Entanglement Entropy
  - Mutual information
- 3. Gravity dual to the CFT

Summary and Future directions

# 2. Global and Local properties of inhomogeneous thermal state

# Thermal entropy



phase transition

 $^{\circ}$  Thermal AdS $_3 \iff au_{
m Mod.} < 1$ 

 $\circ$  BTZ black hole  $\Longleftrightarrow au_{
m Mod.} > 1$ 

• Starting with  $L/\beta < 1$ , the system exhibits the phase transition at

$$au_{ ext{Mod.}}( heta_c) = 1 \Longleftrightarrow heta_c = rac{1}{2} ext{cosh}^{-1}igg(rac{eta}{L}igg) topprox 1_{L/eta \ll 1} rac{1}{2} ext{log}igg(rac{2eta}{L}igg)$$

• In the  $(\chi, \bar{\chi})$  coordinates, the thermal state is given by the usual form with the moduli parameter

$$rac{d_{
m eff}}{eta} = rac{L\cosh(2 heta)}{eta}$$

• From the conventional analysis of the Hawking-Page phase transition, there is a first-order

$$S_{ ext{Thermal}} = egin{cases} \mathcal{O}(c^0) & ext{for } au_{ ext{Mod.}} < 1 \ rac{c \pi L_{ ext{eff}}}{3 eta} & ext{for } au_{ ext{Mod.}} > 1 \end{cases}$$

# Entanglement entropy

for their subsystems

$$egin{aligned} A_1 &= igg\{ x \Big| X_n^f \equiv rac{n}{q} L < X_2 < x < X_1 < X_{n+1}^f igg\} \ A_2 &= igg\{ X_2 < x < X_1 \Big| X_n^f < X_2 < X_{n+1}^f, X_{n+l}^f < X_1 < X_{n+l+1}^f igg\} \end{aligned}$$

$$S_{A_i} = \lim_{n o 1} S_{A_i}^{(n)} = \lim_{n o 1} rac{1}{1-n} \log igg[ rac{\operatorname{tr} \sigma_n(w_1, ar w_1) ar \sigma_n(w_2, ar w_2) e^{-eta H_{q-Mabius}}}{\operatorname{tr} e^{-eta H_{q-Mabius}}} igg] \qquad (w, ar w) = (T+ix, T-ix) = \lim_{n o 1} rac{1}{1-n} \log igg[ \prod_{i=1,2} igg( rac{d\xi_i}{dw_i} igg)^{h_n} igg( rac{d\bar{\xi}_i}{dar w_i} igg)^{ar h_n} igg] + \lim_{n o 1} rac{1}{1-n} \log igg\langle \sigma_n(\xi_1, ar{\xi}_1) ar \sigma_n(\xi_2, ar{\xi}_2) igr\rangle_{\operatorname{Torus} w/ au_{Mod} = A} = rac{L_{ ext{eff}}}{2\pi} \log[\chi], ar{\xi} = rac{L_{ ext{eff}}}{2\pi} \log[ar{\chi}] \quad \chi = igg( rac{\cosh( heta) z^q - \sinh( heta)}{\cosh( heta) - \sinh( heta) z^q} igg)^{ar{q}}, ar{\chi} = igg( rac{\cosh( heta) ar{z}^q - \sinh( heta)}{\cosh( heta) - \sinh( heta) ar{z}^q} igg)^{ar{q}}$$

• Define subsystems  $A_1, A_2$  and compute the entanglement entropies  $S_A = - \text{Tr} \left[ \rho_A \log \rho_A \right]$ 



• To evaluate them, we can use the twist-operator formalism. The Renyi-*n* entropy is given by









#### Entanglement entropy in low effective temperature $au_{ m Mod.} < 1$

In the large  $\theta$ -limit with keeping low effective temperature regime  $L \cosh 2\theta/\beta < 1$ , the entanglement entropy becomes

$$S_{A_1}pprox rac{c}{3} \log \Bigg[ rac{L}{\pi q \epsilon} {
m sin} \Bigg[ rac{q \pi ig( \hat{X}_1 - \hat{X}_2 ig)}{L} \Bigg] \Bigg] \ A_2 pprox rac{c}{3} \log ig( rac{L e^{2 heta}}{2 \pi \epsilon} ig) = rac{2 c}{3} heta + rac{c}{3} \log ig( rac{L}{2 \pi \epsilon} ig)$$

$$S_{A_1}pprox rac{c}{3} {
m log} \Bigg[ rac{L}{\pi q \epsilon} {
m sin} \Bigg[ rac{q \pi ig( \hat{X}_1 - \hat{X}_2 ig)}{L} \Bigg] \Bigg] 
onumber \ S_{A_2} pprox rac{c}{3} {
m log} ig( rac{L e^{2 heta}}{2 \pi \epsilon} ig) = rac{2c}{3} heta + rac{c}{3} {
m log} ig( rac{L}{2 \pi \epsilon} ig)$$

If the subsystem does not include fixed points  $\sin^2(q\pi x/L) = 0$ , the entanglement entropy is given by the vacuum one with system size L/q.



#### Entanglement entropy in high effective temperature $au_{ m Mod.} > 1$

• In the large- $\theta$  limit,

$$\begin{split} S_{A_1} &\approx \frac{c}{6} \log \left[ 4 \prod_{i=1,2} \sin^2 \left( \frac{q \pi X_i}{L} \right) \right] + \frac{c}{3} \log \left( \frac{\beta}{\pi \epsilon} \right) + \frac{c}{3} \log \left[ \sinh \left[ \frac{L \sin \left[ \frac{q \pi (\hat{X}_1 - X_2)}{L} \right]}{2q \beta \sin \left[ \frac{q \pi \hat{X}_1}{L} \right] \sin \left[ \frac{q \pi \hat{X}_2}{L} \right]} \right] \right] \\ &\approx \begin{cases} \frac{c}{3} \log \left[ \frac{L \sin \left[ \frac{q \pi (\hat{X}_1 - \hat{X}_2)}{L} \right]}{q \pi \epsilon} \right] \\ \frac{c}{3} \log \left( \frac{\beta}{\pi \epsilon} \right) + \frac{cL}{6q\beta} \cdot \left[ \frac{\sin \left[ \frac{q \pi (\hat{X}_1 - \hat{X}_2)}{L} \right]}{\sin \left[ \frac{q \pi \hat{X}_1}{L} \right] \sin \left[ \frac{q \pi \hat{X}_2}{L} \right]} \right] \end{cases} \quad \text{For } L \gg \beta \end{split}$$

- If  $L/\beta \ll 1$ , the entanglement entropy reduces to the vacuum one with systems size L/q.

• On the other hand, if  $L/\beta \gg 1$ , it is proportional to  $1/\beta$ , but not simply given by thermal entropy.



• In the large- $\theta$  limit,

$$S_{A_2} pprox rac{c}{6} \log \Biggl[ 4 \prod_{i=1,2} \sin^2 \Biggl( rac{q \pi X_i}{L} \Biggr) \Biggr] + rac{c}{3} \log \Biggl( rac{eta}{\pi \epsilon} \Biggr) + rac{c \pi L e^{2 heta}}{6eta} \cdot rac{l}{q} pprox rac{c}{3} \log \Biggl( rac{eta}{\pi \epsilon} \Biggr) + rac{c \pi L e^{2 heta}}{6eta} \cdot rac{l}{q}$$

 The entanglement entropy is proportional to number of fixed points size

 $L_{\rm eff}$ 

• Thermal contributions are almost localized at fixed points

 $sin^2(q\pi x/L) = 0$  included in the subsystem  $A_2$ , i.e. l, and to the effective system

$$pprox L \cdot rac{e^{2 heta}}{2}$$



q = 4

#### First-order phase transition of Entanglement Entropy • Starting with $L/\beta < 1$ and increasing $\theta$ , we can see both high and low effective temperature regions of the entanglement entropies for the two regions $A_2$ $A_1$ **Entropy Gap** No Entropy Gap $A_{2}$ $1.2 \times 10^{8}$ 13.7230 $1.0 \times 10^{8}$ 15.0 13.7228 13.7226 14.8 $8.0 \times 10^{7}$ 13.7224 S 14.6 13.7222 $5^{-107} 6.0 \times 10^{7}$ 14.4 14.2 3.8 3.9 3.6 3.7 4.0 $4.0 \times 10^{7}$ 2.0 2.2 2.4 2.6 2.8 3.0 $2.0 \times 10^{7}$



Н  $q = 4, L/\beta = 10^{-2}, L = 100000,$  $X_1 = 2L/q + 10^3 + 10, X_2 = 10^3 + 10, c = 1, \varepsilon = 10^{-14}$ 

8

10

12



# Mutual information

- The enveloping function is related to a local temperature.
- Let us see the relation by considering the mutual information between subsystems,

$$I(A_1, A_1') = S_{A_1} + S_{A_1'} - S_{A_1 \cup A_1'}$$
  
 $A_1 = \left\{ x \Big| X_n^f \equiv \frac{n}{q} L < X_2 < x < X_1 < X_{n+1}^f \right\} \quad A_1' = \left\{ x \Big| X_n^f < X_1 < X_4 < x < X_3 < X_n^f \right\}$   
• If we focus on the uniform case with  $L/\beta \gg 1, \theta = 0$ , the mutual information is just given by

$$Iig(A_1,A_1'ig)ig|_{L/eta\gg 1,{
m w}/ heta=0}pprox \maxigg\{0,-rac{c}{3} \logigg| \expigg[rac{2\pi}{eta}(X_4-X_1)igg]-1igg]igg\},$$

implying the correlation length (defined by the distance that the mutual information becomes 0) related to the inverse temperature

$$d_{41, heta=0}=rac{eta}{2\pi}{\log 2}$$



Åγ

regions

$$au_{
m Mod.} = L_{
m eff}/eta, L/eta \gg 1, heta \gg 1$$

• Then, the mutual information can be evaluated as

 $\left. I(A_1,A_1') 
ight|_{L/eta \gg 1 ext{ w}/ heta \gg 1}$ 

$$pprox \max iggl\{ 0, -rac{c}{3} \log \left[ \exp \left[ rac{2L}{qeta} \cdot rac{\sin \left( rac{q\pi}{L} \left( \hat{X}_4 - \hat{X}_1 
ight) 
ight)}{\cos \left( rac{q\pi}{L} \left( \hat{X}_4 - \hat{X}_1 
ight) 
ight) - \cos \left( rac{q\pi}{L} \left( \hat{X}_1 + \hat{X}_4 
ight) 
ight)} 
ight] - 1 
ight] iggr\}$$

This implies the correlation length related to the "effective local temperature" depending on position

$$d_{41, heta\gg1}pprox rac{eta}{2\pi}\cdot f\!\left(rac{\left(\hat{X}_1+\hat{X}_4
ight)}{2}, heta
ightarrow\infty
ight)\log 2$$

• To find a similar (local) temperature for the inhomogeneous case, we focus on the parameter



$$figg(rac{\left(\hat{X}_1+\hat{X}_4
ight)}{2}, heta
ightarrow\inftyigg)=1-\cos\Bigl(rac{q\pi}{L}\Bigl(\hat{X}_1+rac{q\pi}{L}\Bigr)$$



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Summary and Future directions

# 3. Gravity dual to the CFT

#### Gravity dual to the inhomogeneous thermal state

- by the usual form.
- given by the standard BTZ black hole [Banados-Teitelboim-Zanelli '92]

$$ds^2_{
m bulk} \, = rac{dr^2}{r^2 - r_0^2} - r^2 igg( rac{d \xi_ au - d ar{\xi}_ au}{2} igg)^2 + ig( r^2 - r_0^2 ig) igg( rac{d \xi_ au + d ar{\xi}_ au}{2} igg)^2$$

$$(\xi_{ au}, ar{\xi}_{ au}) = (\xi(z) + au, ar{\xi}(ar{z}) + au) \qquad r_0 = rac{2\pi}{eta}$$

• In the original coordinates,

$$egin{aligned} ds^2_{ ext{bulk}} &= rac{dr^2}{r^2 - r_0^2} + ig(r^2 - r_0^2ig) d au^2 + rac{r^2}{f(X, heta)^2} dX^2 \ f(x, heta) &= 1 - anh 2 heta ig(1 - 2 \sin^2ig(rac{q\pi x}{L}ig)ig) &= 1 - anh 2 heta \cdot \cosig(rac{2q\pi x}{L}ig) \end{aligned}$$

$$egin{aligned} s_{ ext{bulk}}^2 &= rac{dr^2}{r^2 - r_0^2} + ig(r^2 - r_0^2ig) d au^2 + rac{r^2}{f(X, heta)^2} dX^2 \ f(x, heta) &= 1 - anh 2 heta ig(1 - 2 \sin^2ig(rac{q\pi x}{L}ig)ig) &= 1 - anh 2 heta \cdot \cosig(rac{2q\pi x}{L}ig) \end{aligned}$$

• In the  $(\chi, \bar{\chi})$  or  $(\xi, \bar{\xi})$  coordinates, the thermal state for the inhomogeneous Hamiltonian is given

• Then, the gravity dual to the thermal state for high effective temperature  $L \cosh(2\theta)/\beta > 1$  is



• In the large-*r* region,

$$ds^2_{
m bulk} \left|_{r \gg 1} pprox rac{dr^2}{r^2} + rac{r^2}{f(X, heta)^2} ig\{ f(X, heta)^2 d au^2 + dX^2 ig\}.$$

This include the (Euclidean) CFT background metric.

• Or, by introducing the new radial coordinate

$$r'(X, heta) = rac{r}{f(X, heta)}$$

then, we obtain the standard asymptotic metric

$$ds^2_{
m bulk} \left|_{r' \gg 1} pprox rac{dr'^2}{r'^2} + r'^2 ig\{ f(X, heta)^2 d au^2 + dX^2 ig\}$$

dependent



In terms of the new radial coordinates, the horizon radius becomes position-

# Effective local temperature

- In the new radial coordinates, we can estimate an effective local temperature.  $\bullet$
- ullettemperature

$$ilde{eta}_{ ext{eff},x=X; heta}=rac{2\pi}{r_0'(X, heta)}=rac{2\pi}{r_0}f(X, heta)=eta f(X, heta)$$

Let us assume an observer sitting at X; the typical scale of the observer is assumed to be comparable with the re-scaled horizon radius. Such an observer feels the local effective





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Summary and Future directions

## **Summary and Future directions**

- We studied the thermal state with the CFT Hamiltonian on the curved space
  - The CFT Hamiltonian on the curved space is conformally equivalent to that on the flat space with size  $L \cosh(2\theta)$
  - By evaluating the thermal entropy, we found the first-order phase transition by varying the background metric using  $\theta$  with fixing  $L/\beta < 1$ .
  - If the subsystem includes fixed points  $\sin^2(q\pi x/L) = 0$ , the entanglement entropy exhibit the first-order phase transition, but if not, it does not exhibit the phase transition.
- Related works and future directions
  - Entanglement dynamics (using the quantum quench and the state-operator mapping) associated with the CFT  $\bullet$ Hamiltonian [Bai-Miyata-Nozaki arXiv: 2408.06594 [hep-th]]
  - Bulk reconstruction and relative entropy?
  - Quantum correlation of 2d Holographic CFTs on the curved space

• • •

Thank you!!













